

# Dynamic Sky Dome

KGC November, 2011

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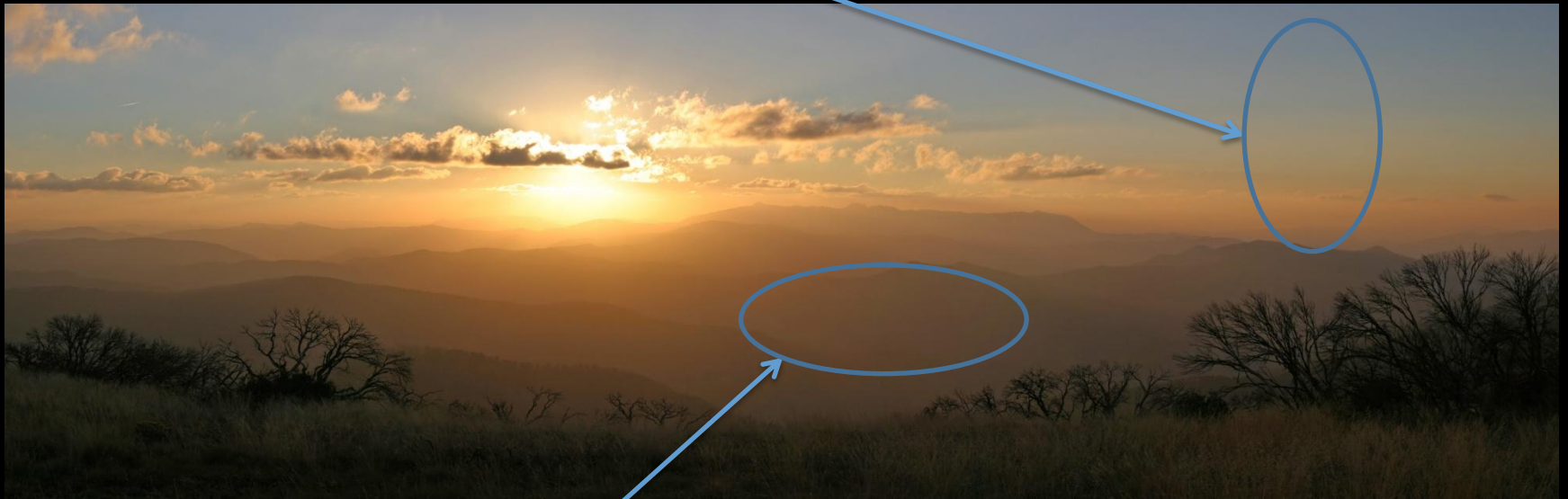


# Agenda

- Dynamic Sky Dome lighting model
  - Introducing the rendering equation
  - Fog integration
  - Sun shafts
- Dynamic clouds rendering
- Bonus trick

# Effects to Achieve

- Rayley scattering
- High chromatic dependency



- Mie scattering
- Achromatic

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# Sky Dome Rendering History

- [PSS99][HP03][O'N05]
- Single scattering only
- Or viewable from ground only
- Or all of the above
- Or not real-time

# Current Approach

- Multiple scattering
- Aerial perspective
- Viewable from space
- LUTs are pre-computed on GPU

# Rendering Equation

$$L(x, v, s) = (L_0 + R[L] + S[L])(x, v, s)$$

- $x$  – position,  $v$  – view direction,  $s$  – sun direction
- Final lighting includes:
  - $L_0$  – attenuated direct sun light
  - $R[L]$  – light reflected at point  $(x_0)$  and attenuated
  - $S[L]$  – light scattered toward the viewer between  $x$  and  $x_0$
- Usually computed by numerical integration

# Direct Sun Light

$$L_0(x, v, s) = T(x, x_0) L_{\text{sun}}, \text{ or } 0$$

- Direct sunlight attenuated by the transmittance function  $T(x, x_0)$  before reaching viewer at  $x$
- Accounts for occlusion by the horizon
- Smooth shadow transition to hide sharp border on the clouds

# Reflected Light

$$R[L](x, \nu, s) = T(x, x_0) I[L](x_0, s)$$

- Light reflected at  $x_0$  and attenuated by transmittance towards  $x$
- $I[L] = 0$  at the top atmosphere boundary

# Inscattered Light

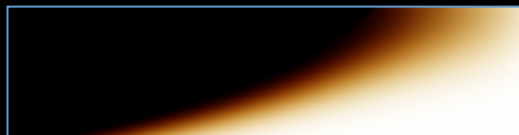
$$S[L](x, v, s) = \int_x^{x_0} T(x, y) J[L](y, v, s) dy$$

- Light scattered between  $x$  and  $x_0$  toward  $x$

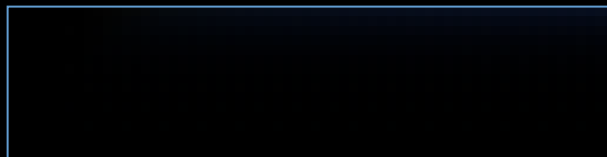
# Look-up Tables (LUT)

- Re-parametrise functions and store them in LUT
- Pre-compute LUTs on GPU

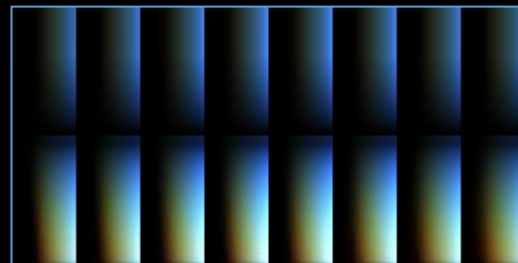
Transmittance ( $r, \mu$ )



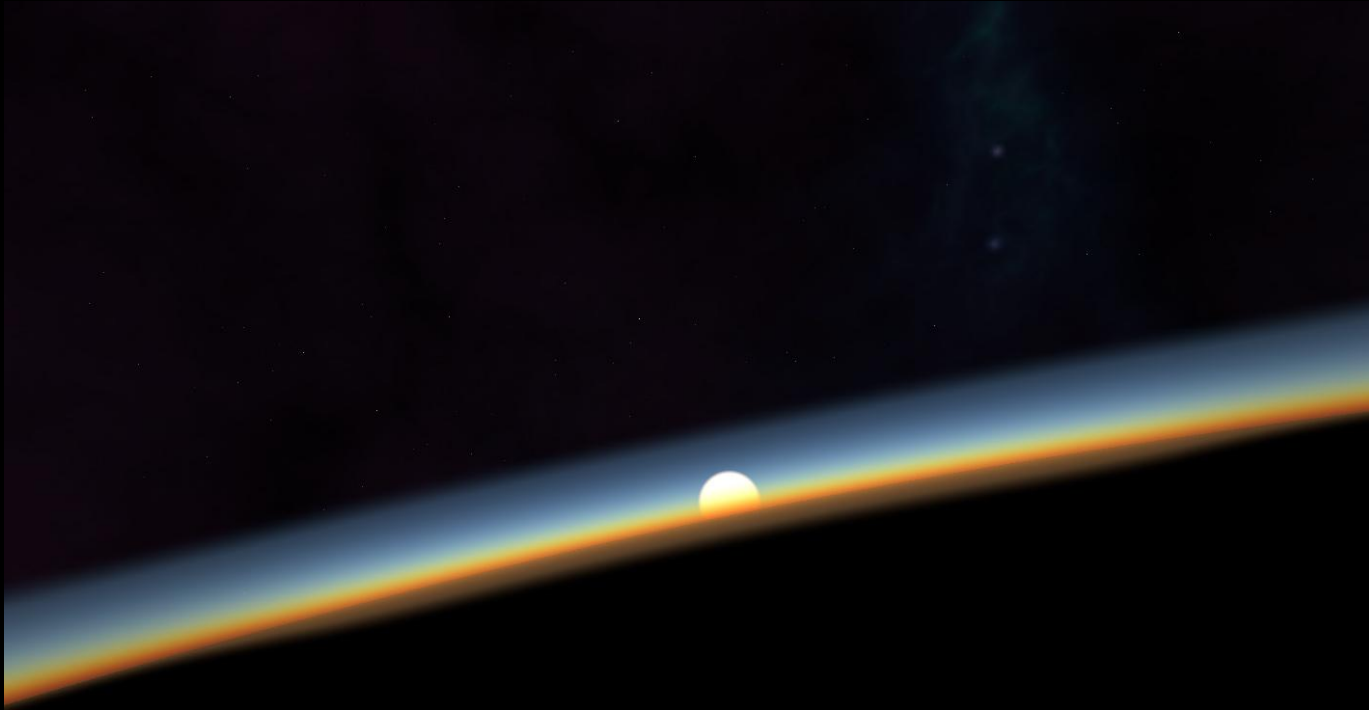
Irradiance ( $r, \mu S$ )



Inscatter ( $r, \mu, \mu S, \nu$ )



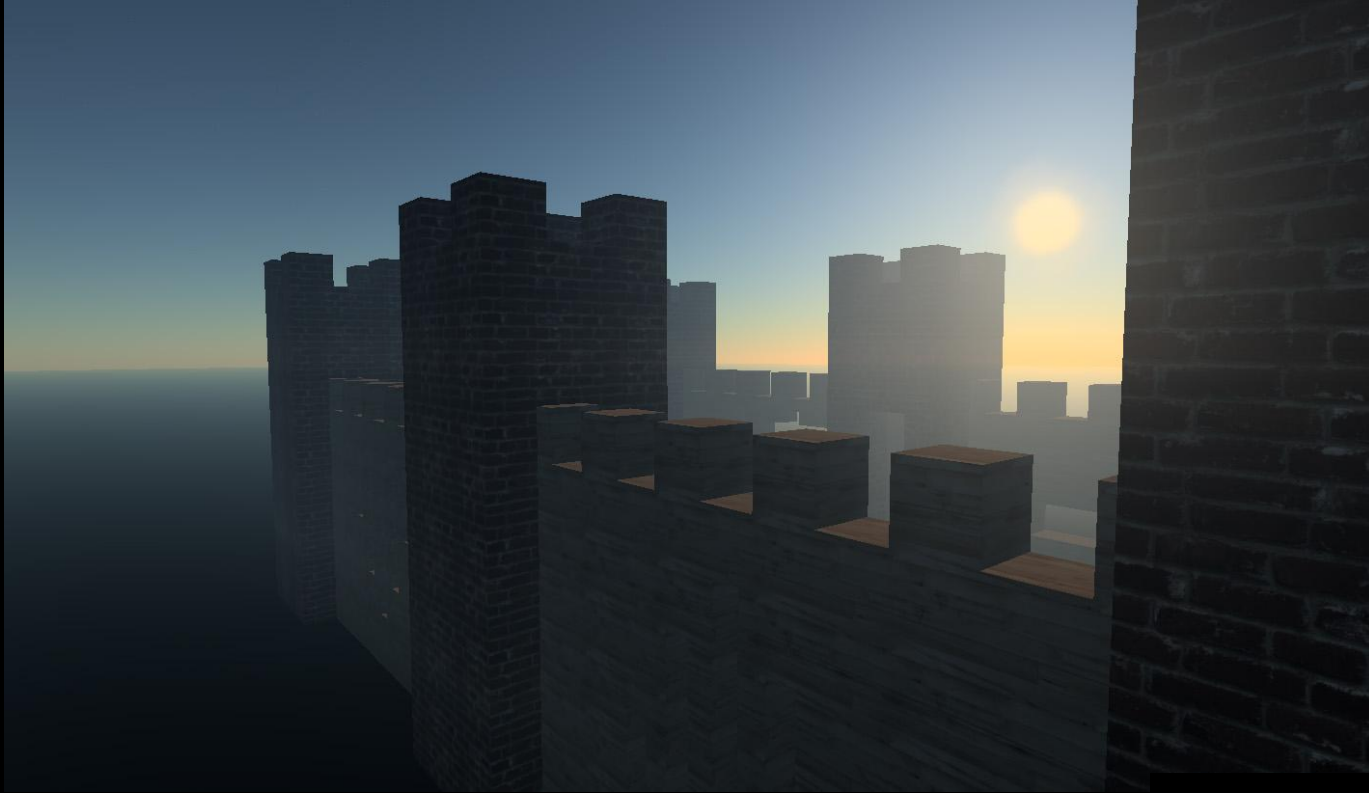
# View from Space



# Simple Fog vs Inscatter

- Fast
  - Change parameters in real time
  - Accounts for transmittance only
  - Looks fine
- Slower
  - Precompute
  - Accounts for transmittance and inscatter
  - Extremely realistic

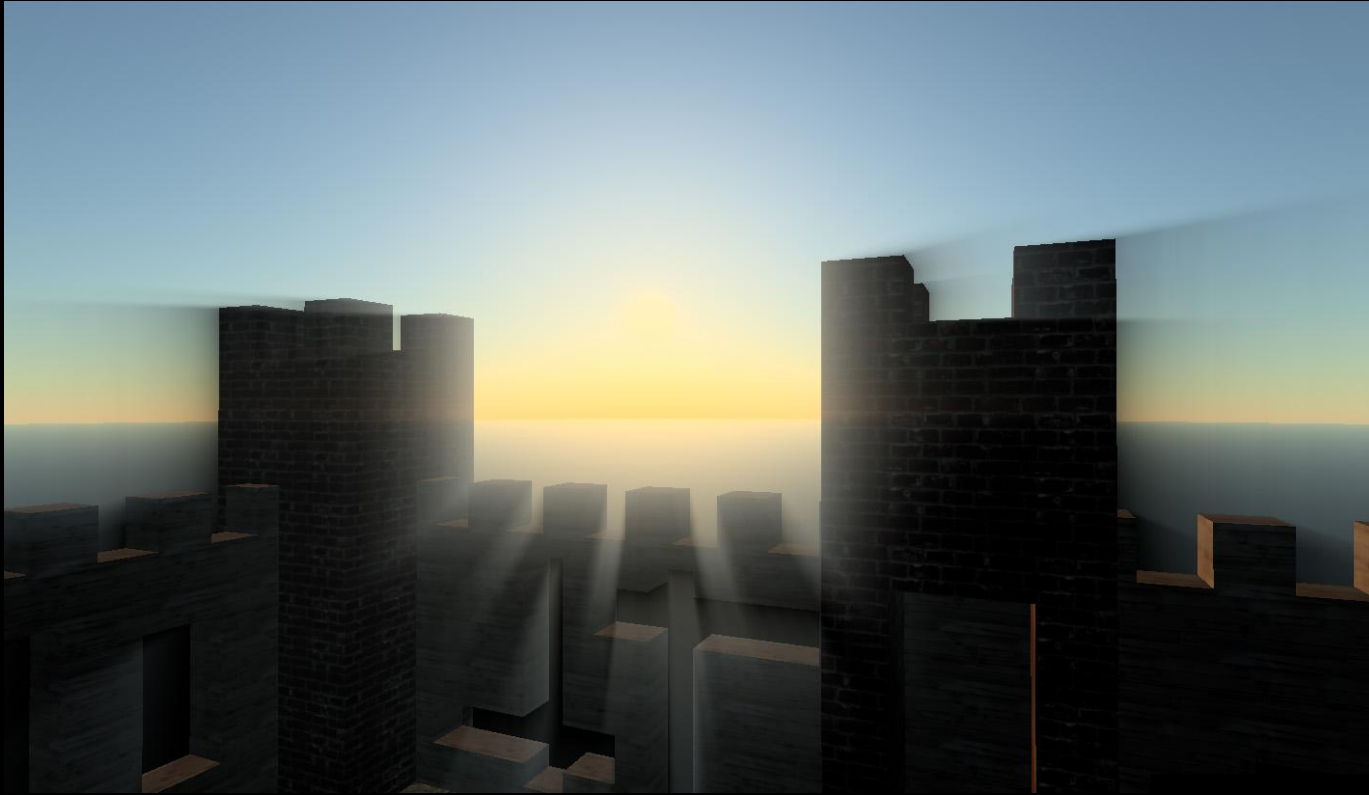
# Inscatter



# Sun shafts

- Use industry-standard mask generation
- Radial blur it
- Modulate in-scattered light with the mask
- Cheap, known, works for all types of geometry

# Sun shafts



# Clouds

- Need volumetric realistically looking clouds
- Use group of particles to represent a single cloud
- Use the same lighting equation to blend nicely into the sky
- Use impostors to save on expensive per-pixel lighting computations

# Clouds lighting

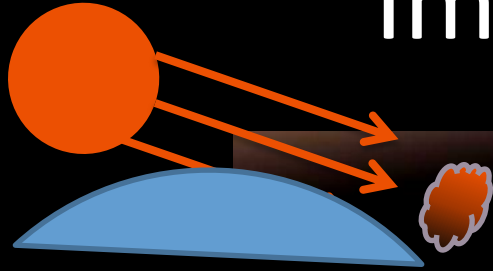
- Need  $x_0$  to apply lighting equation
  - Use world-space impostor's pixel position
- Apply inscattered light and irradiance as to a regular object.
- To simulate sun direct light scattering use similar to [WENZEL07] technique
  - March in sun direction in impostor space to accumulate light extinction factor to attenuate sun light
  - Add rim light to simulate scattering on the cloud's edge

# Clouds lighting

143 fps (7.48 ms)



# Impostor Pixel Position



# Impostor Pixel Position

- Need better pixel position reconstruction:
  - Sun shadow is a crisp horizontal line
  - Can't pack a group of clouds into a single impostor or have a very large cloud

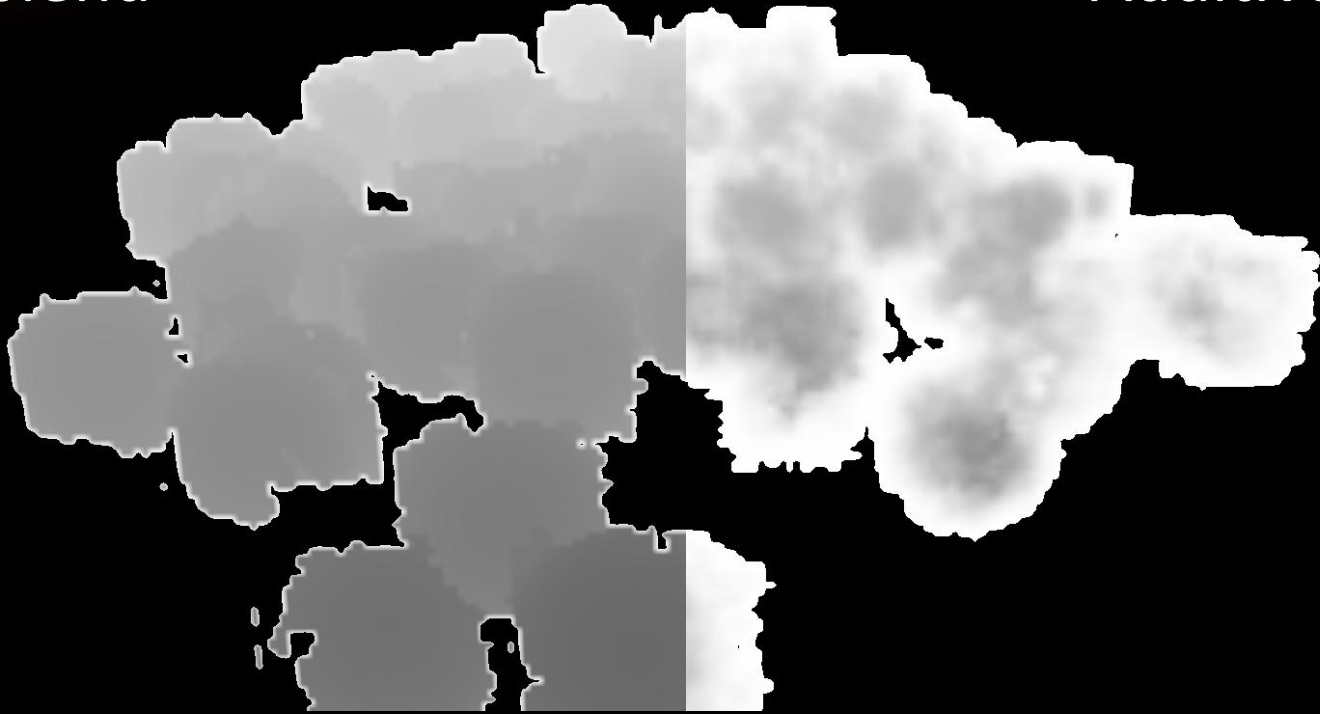
# Impostor Pixel Position

- Accumulate impostor-space linear depth per pixel
  - Simple min doesn't work well – must account for transparency, linear interpolation problem on borders
  - Simple additive blend doesn't work – depth on cloud edges blend with impostor clear depth value

# Impostor Pixel Position

Min blend

Additive blend



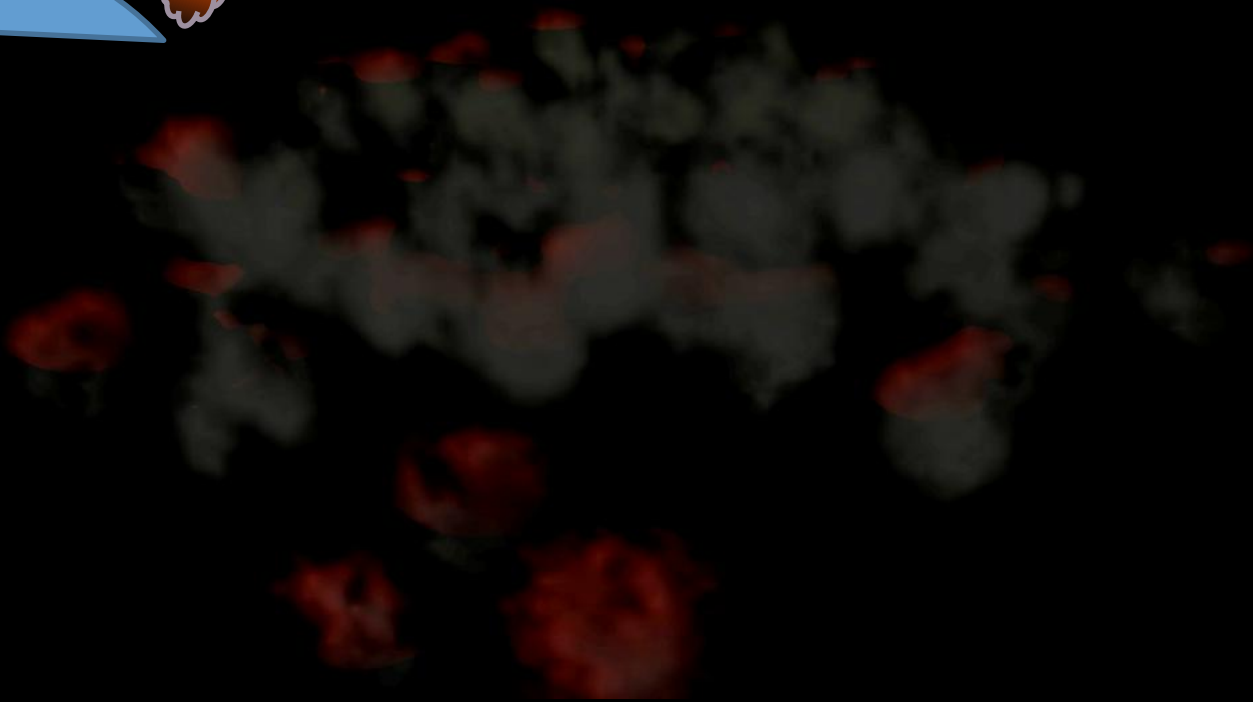
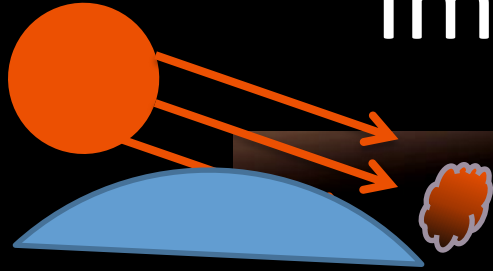
# Impostor Pixel Position

- Store both accumulated depth and normalization weight
  - $\text{DepthAcc} = \text{OldDepthAcc} * (1 - \text{newDepthWeight}) + \text{NewDepth} * \text{NewDepthWeight}$
  - $\text{DepthWeightAcc} = \text{OldDepthWeightAcc} * (1 - \text{newDepthWeight}) + \text{NewDepthWeight}$
  - Final depth =  $\text{DepthAcc} / \text{DepthWeightAcc}$
- Use particle transparency as DepthWeight => Need 2 FP16 channels to store both impostor depth and density
- Linear interpolation is not linear anymore, but it is monotonous and works
- Have two more spare channels – store normal or per-particle ambient occlusion for more sophisticated lighting?

# Impostor Pixel Position



# Impostor Pixel Position

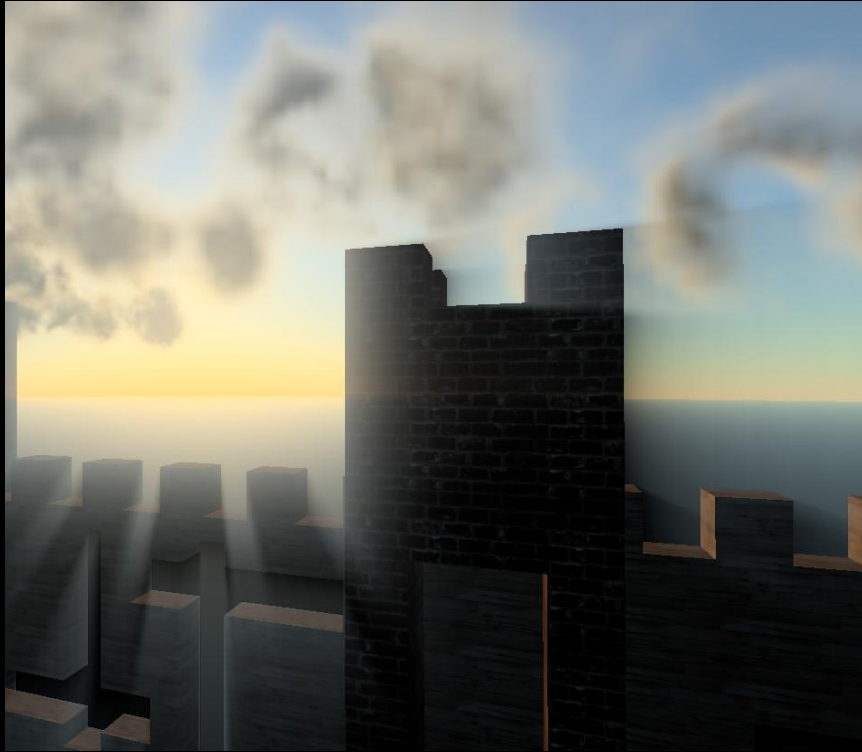


# Polarization

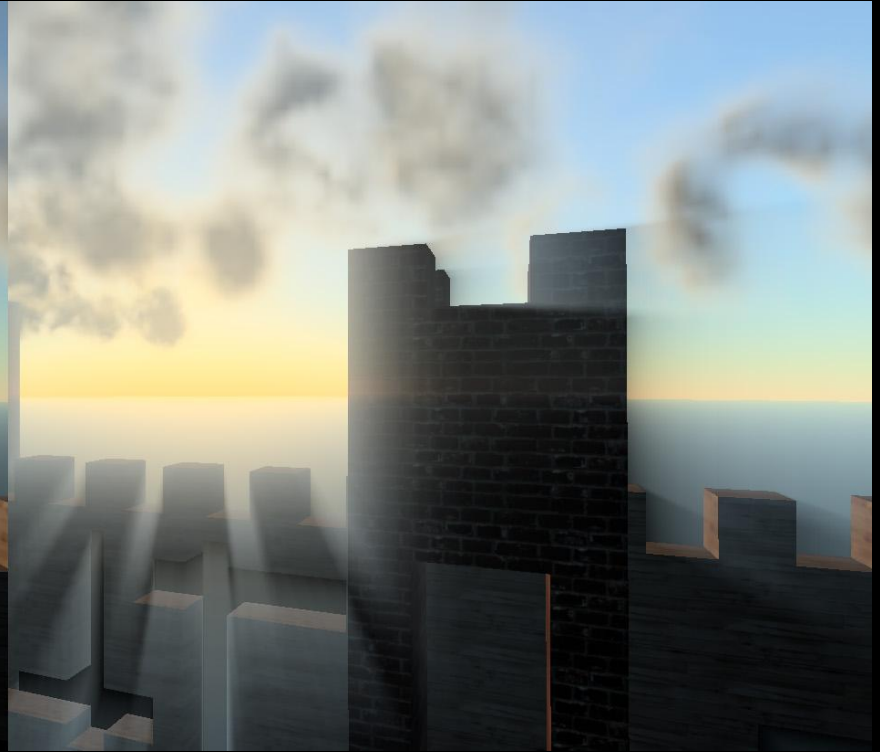
- Sky light is polarized
  - Max polarization at view-sun angle ==  $90^{\circ}$
  - Min polarization at view-sun angle ==  $0^{\circ}$  and  $180^{\circ}$
- Simulate effect of camera with polarized filter attached to add more contrast and depth to the scene

# Polarization

ON



OFF



# Questions ?



# Reference

- [PSS99] Preetham, A. J., Shirley, P., and Smits, B. E.: A Practical Analytic Model for Daylight. In *Siggraph 1999, Computer Graphics Proceedings* (Los Angeles, 1999), Rockwood A., (Ed.), Addison Wesley Longman, pp. 91–100.
- [HP03] Hoffman, N., and Preetham, A. J.: Real-time Light-Atmosphere Interactions for outdoor scenes. *Graphics Programming Methods* (2003), pp. 337–352.
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